

Astrophysics I: Stars and Stellar Evolution

AST 4001

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Stars and Stellar Evolution, Fall 2008

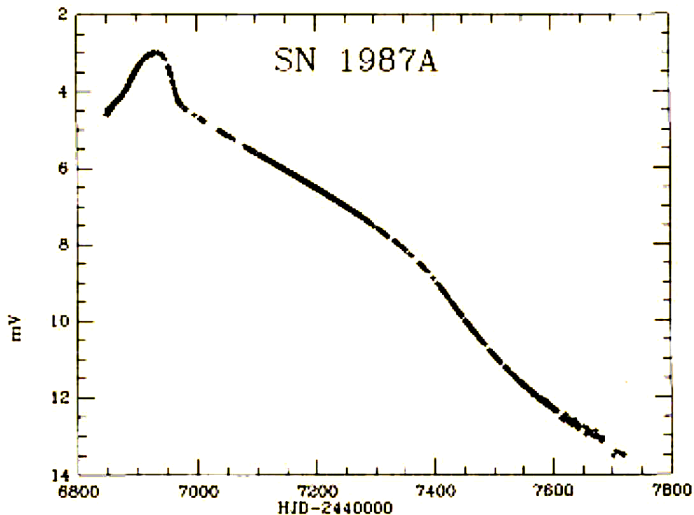
Overview

- 1 Recap
- 2 Supernovae
 - Supernovae - Observations
 - SN 1987A
 - Supernova Remnants
- 3 Supernova Summary
 - Remnants as a Function of Mass
 - Supernovae as a Function of Mass and Metallicity
 - Remnants as a Function of Mass and Metallicity

SN 1987A



SN 1987A Visual Light Curve



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Core Collapse Supernovae - Observations

Type Ib/Ic/II

- $\sim 75\%$ of all supernovae
(the rest are “Type Ia”)
- only occur in regions near young stars, molecular clouds \Rightarrow
originate from short-lived (i.e., most massive) stars
- neutron stars found associated with historical Type Ib/Ic/II
- now accepted that these are all “core-collapse” supernovae

Core Collapse Supernovae - Observations

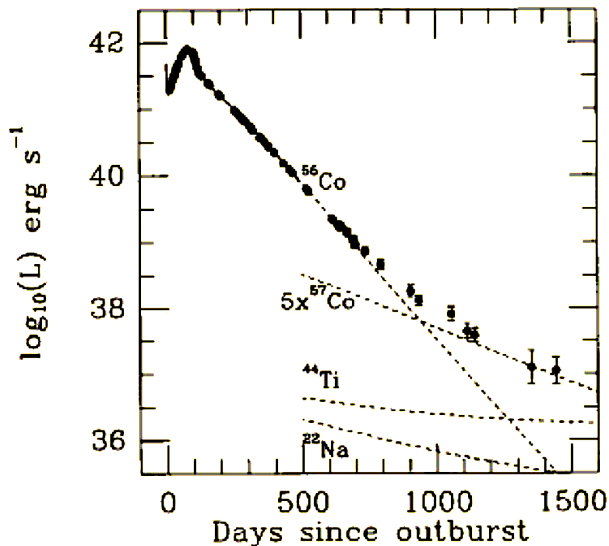
Different progenitor stars result in supernova subclasses

- II-P: normal isolated $10 - 30 M_{\odot}$ star
- II-L: as for II-P, but with smaller hydrogen envelope?
- Ib: very massive star; H envelope lost through wind or binary
- Ic: as for Ib, but He envelope lost also? Hypernovae?
- IIn: explosion expands through dense clumpy wind?

Core Collapse Supernovae - Observations

- Wide variation in brightness: $L_{\text{peak}} \sim (0.4 - 4) \times 10^9 L_{\odot}$
- After peak, light curve dominated by radioactive decay
- new elements are unstable
- decay produces photons, e.g.,
 - $^{56}\text{Ni}(e^+\nu_e)^{56}\text{Co}$ (average life: 8.8 days)
 - $^{56}\text{Co}(e^+\nu_e)^{56}\text{Fe}$ (111 days)
 - $^{57}\text{Co}(e^+\nu_e)^{57}\text{Fe}$ (391 days)
 - $^{44}\text{Ti}(e^+\nu_e)^{44}\text{Sc}$ (87 years)
- mechanism underlying slow fade of “light curve”

SN 1987A - Radioactive Decay



SN 1987A - Facts

- 23 Feb 1987: new SN in Large Magellanic Cloud ($d = 50$ kpc)
- Peak magnitude $V \sim 3$: first naked eye SN since 1604!
- 19 neutrinos detected! Energy, composition, time, direction all consistent with formation of a neutron star
- Progenitor star was Sk -69°202, a (blue) supergiant
- No neutron star / black hole detected yet
- Light curve clearly shows phases of radioactive decay

Supernova Energetics

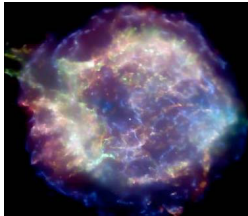
- Core collapse Supernovae (neutron star - Type Ib/c, II)
 - $\sim 3 \times 10^{53}$ erg in neutrinos
 - $\sim 10^{51}$ erg in mechanical (kinetic) energy
 - $\sim 10^{49}$ erg in (visible) photons
 - $\sim 0.1 M_{\odot} {}^{56}\text{Ni}$
- Type Ia supernovae (thermonuclear)
 - $\sim 10^{51}$ erg in mechanical (kinetic) energy
 - $\sim (2 - 3) \times 10^{49}$ erg in (visible) photons
 - $\sim 0.5 M_{\odot} {}^{56}\text{Ni}$

origin of ${}^{56}\text{Fe}$ in the sun:

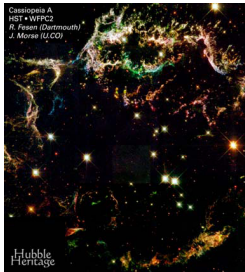
- $\sim \frac{1}{3}$ from core collapse supernovae
- $\sim \frac{2}{3}$ from Type Ia supernovae

Supernova Remnants

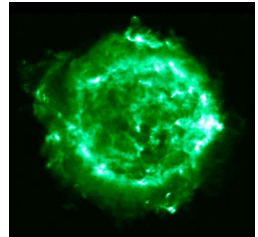
- Radioactivity fades in 50-100 years
- Explosion drives shock wave into surrounding gas supernova remnant (SNR)
- Shock acceleration: synchrotron emission (radio)
- Shock heating: bremsstrahlung (X-rays)
- Collisions with clumps and clouds: emission lines (optical)



Cas A: X-ray



optical



radio

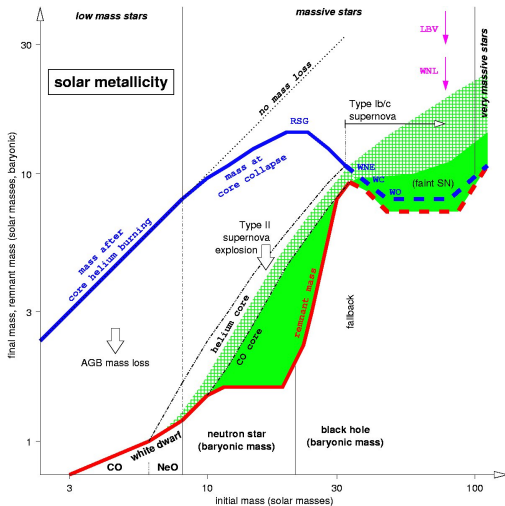
Supernova Remnant Evolution

- “Free expansion”: ejecta expand freely with $E_{\text{SN}} = \frac{1}{2} M_{\text{ejecta}} v^2$
- ”Adiabatic” or “Sedov-Taylor” phase
 - begins when mass swept up $>$ mass ejected
 - energy conserved (very little escapes in radiation)
 - 50 % of energy is kinetic, 50 % of energy is thermal
 - $R = \left(\frac{25}{3\pi}\right)^{1/5} \left(\frac{E_{\text{SN}}}{\rho}\right)^{1/5} t^{2/5}$
- “Radiative” or “snowplow” phase
 - begins when 25 % of total energy lost as radiation (typically when $v \sim 200$ km/s)
 - expansion now driven by conservation of momentum
 - evolution: $R \propto t^{1/4}$
- “Dissipation” - SNR stops expanding and becomes static, hot, bubble when $v < c_{\text{sound}} \Rightarrow t_{\text{fade}} \sim 3$ Myr, $R_{\text{fade}} \sim 75$ pc

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Stellar Mass Ranges - Solar Metallicity



Stellar Mass Ranges - Population III Stars

